Fundamentals – Radiation Interactions



- Types of radiation and means of energy conversion
- Interaction/ detection processes
 - Charged particles
 - Electrons
 - Photons
 - Neutrons



General Principles of Radiation Detection

- Radiation detection
 - Interaction of radiation with matter produces ionization and electronic excitation or heat that can be measured:
 - Either primary charges are collected:
 - Gas detectors
 - Solid state detectors

 Or photons resulting from deexcitation of molecules of the detector are converted to secondary charges which are collected:

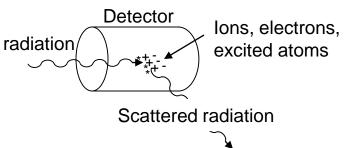
• Scintillators

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Ionization chamber Proportional counter Geiger-Müller counter

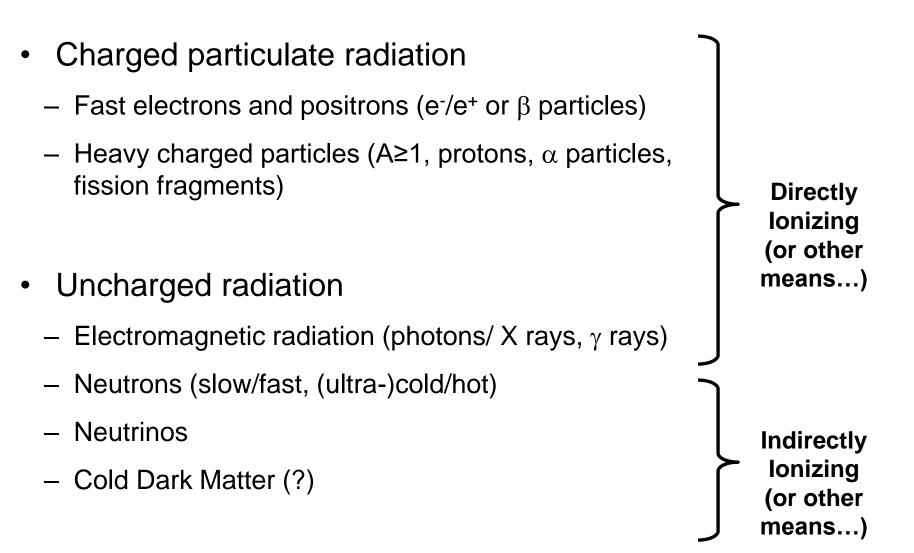
Si, Ge, CdZnTe, Hgl₂,...

Inorganic: Nal(TI), Csl(TI), LaBr, BGO,... Organic: anthracence, stilbene, plastic,...









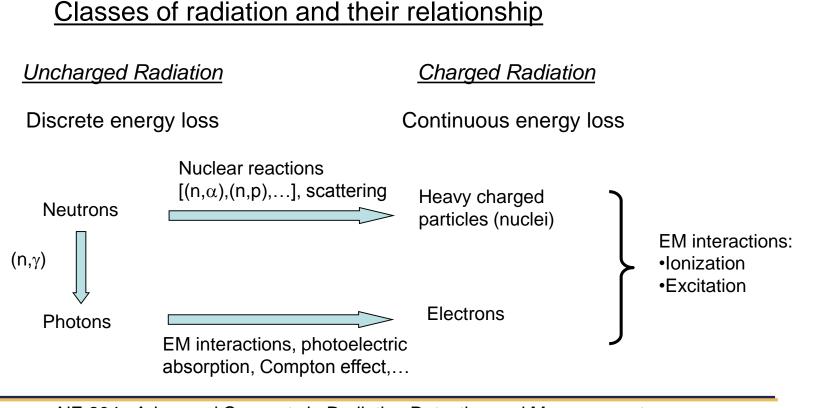
Some Properties of Ionizing Radiation



Heavy charged particles	Energy when Generated			
 α-decay 	Discrete			
Spontaneous fission	Continuous			
Electromagnetic radiation				
 Gamma rays following beta decay or other means of nuclear excitation 	Discrete			
Annihilation radiation (511 keV)	Discrete			
Bremsstrahlung	Continuous			
Characteristic X rays	Discrete			
Neutrons				
 Spontaneous and induced fission 	Continuous			
 Radioisotope (α,n) sources 	Continuous			
 Photo-neutron (γ,n) sources 	~ Discrete			
 Accelerated-based neutron generators [(D,D); (D,T); (p/d,n) reactions] 	~ Discrete			



 To understand radiation detection, it is necessary to understand underlying physics processes how radiation interacts with matter, e.g. detectors...



Means of and materials for converting energy to signal



• Ionization, Scintillation, Heat vs. Gases, Liquids, Solids

Material	Detector	Signal	Excitation
State	implementation		energy
Gas	Scintillation	Light - Photons	10-200 eV
	Ionization	Electron-ion pairs	~ 30 eV
Liquid	Scintillation	Light - Photons	10-200 eV
	Ionization	Electron-ion pairs	~ 30 eV
Solid	Scintillation	Light - Photons	10-200 eV
	Ionization	Electron-hole pairs	1-5 eV
	Bolometer	Heat - Phonons	~ 0.001 eV

- And combinations of implementations, e.g.
 - Gas & liquid: Scintillation (prompt) + ionization (delayed): Particle discrimination (nuclear vs. electronic), energy resolution improvements, 3D position determination (Time-Projection Chamber)
 - Solid: Ionization + Bolometer: Particle discrimination (nuclear vs. electronic)

Other means of detection Even non-EM radiation



- Detect by different interaction process as a way to distinguish particle types to increase sensitivity by recognizing background ... important in the detection of rare particles and processes such as CDM or v's ...
 - For example: CoGent, Majorana Ionization Edelmer's Nuclear Recoil, OvBB, etc Scintillation teraction, CRESST III, Rosebud

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7

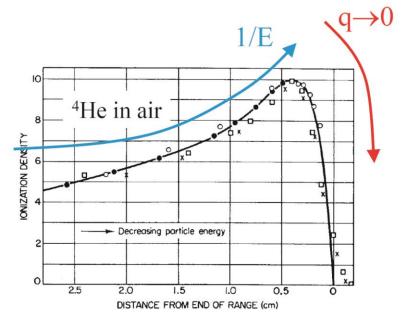
Review of Interactions

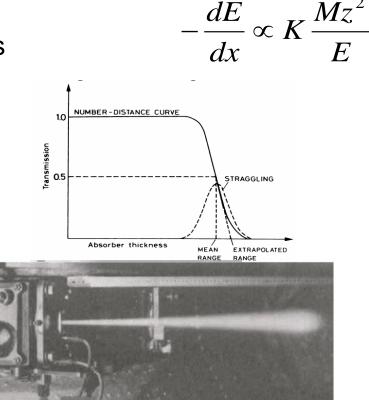


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Interaction of Massive Charged Particles

- Charged particles experience energy loss and deflection due to interaction with:
 - -Inelastic collisions with atomic electrons
 - -Elastic scattering on nuclei
 - -Bremsstrahlung

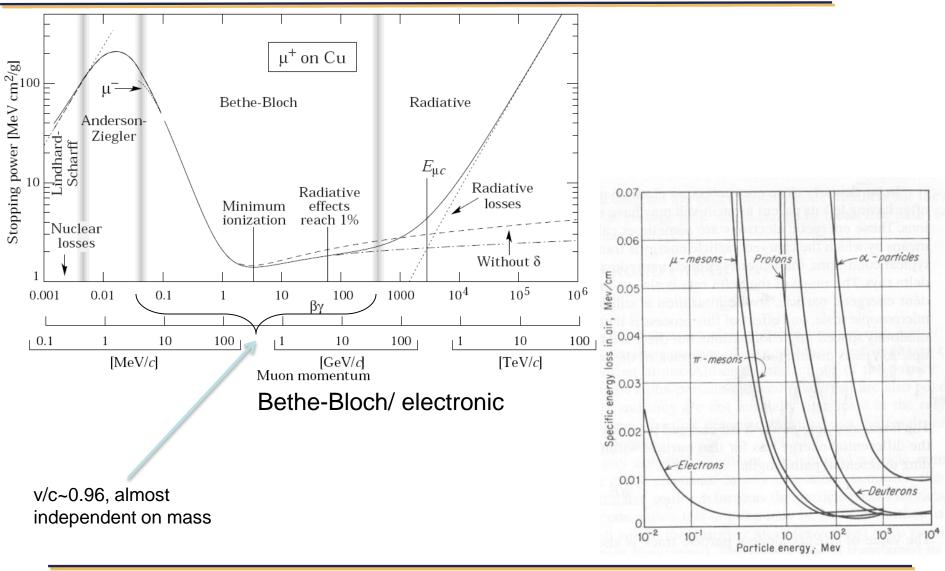




Deuterons in air from: A.K. Solomon, "Why Smash Atoms?" (1959)



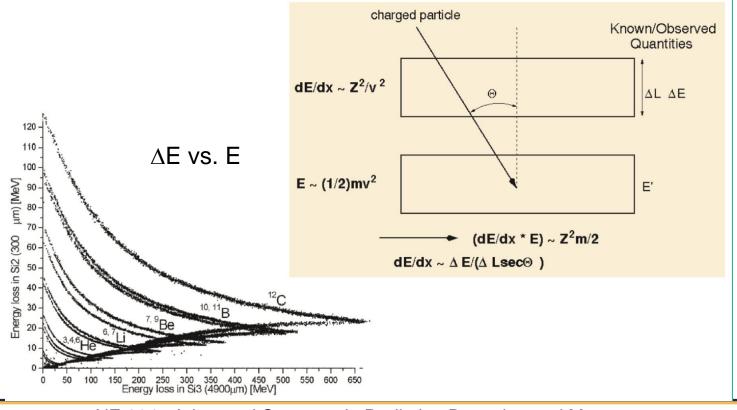
Stopping Power







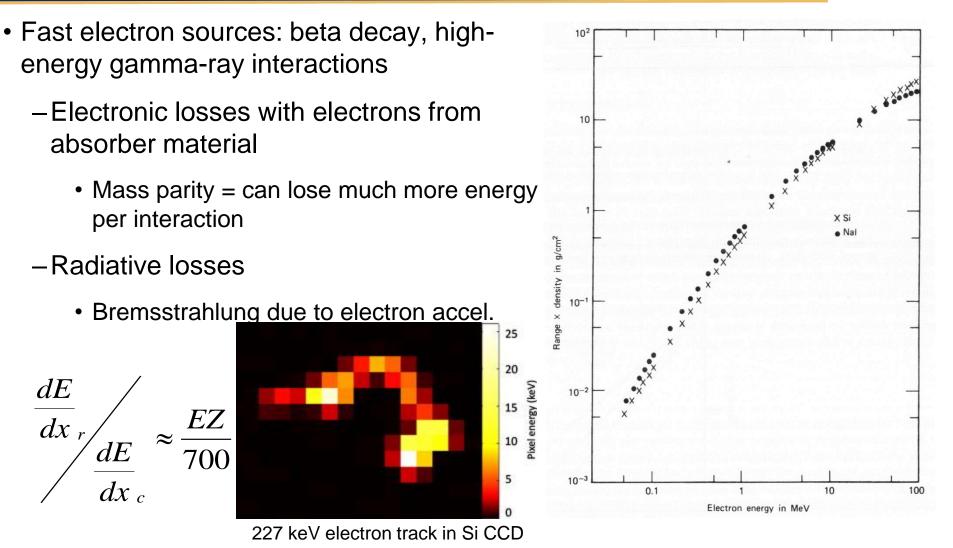
- E.g. ⁴⁸Ca + ²⁰⁸Pb @ 200 MeV (P. Reiter, T.K. Khoo, Argonne National Laboratory):
 - Reaction products identification with ΔE -E telescope:



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Interaction of Fast Electrons





Interactions of Photons/ Gamma Rays



- A beam of photons passes through material until each undergoes a collision at random and is removed from beam
 - Intensity continuously drops, but energy remains constant (in contrast to heavy charged particles which slow down continuously without losing intensity)

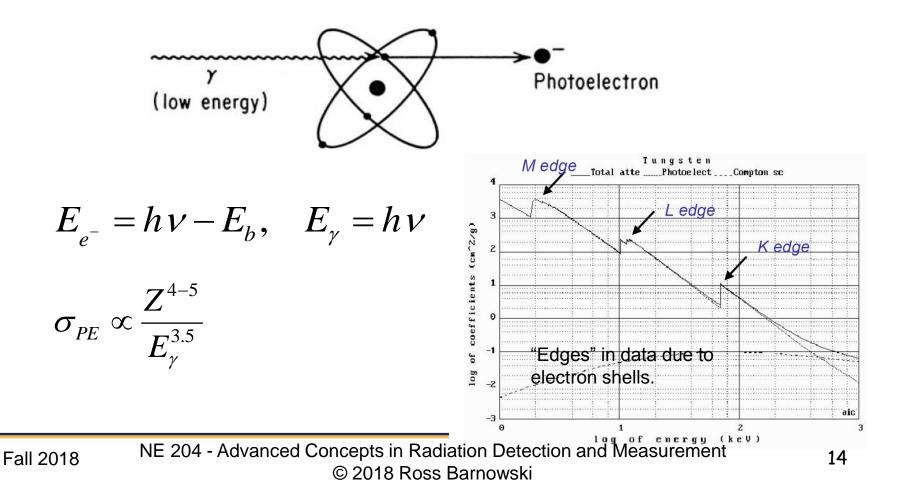
$$I = I_0 e^{-\mu x}, \quad \mu = \frac{1}{\lambda}$$

 $\label{eq:main_linear} \begin{array}{l} \mu \text{: attenuation coefficient} \\ \lambda \text{: mean free path} \end{array}$

- Four interaction processes:
 - Photoelectric absorption
 - Compton scattering
 - Pair production
 - Coherent or Rayleigh scattering (elastic)

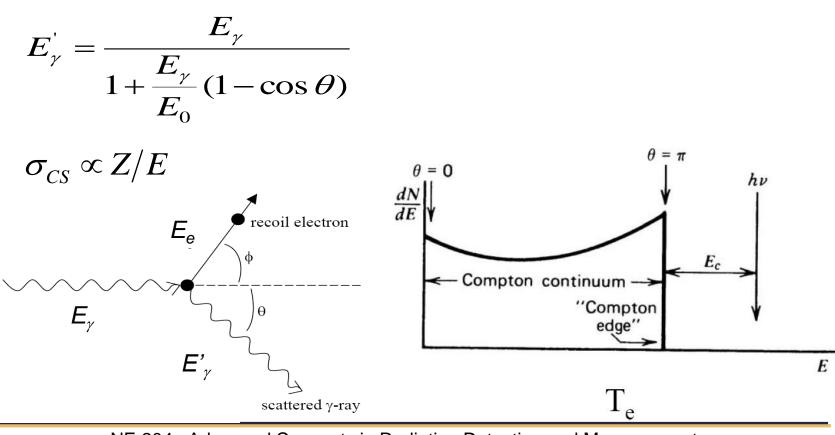


 Entire photon energy is transferred to a bound (most likely K-) electron:





 Scattering of a photon by a (free) electron that leads to a moving electron and a lower energy photon:



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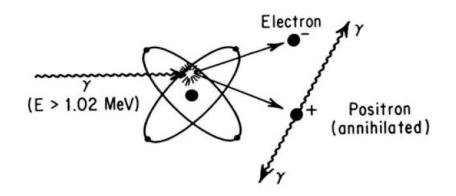
Pair Production



- For E_g>1.022 MeV, the photon can be converted into an electron-positron pair in the presence of a nucleus.
- After slowing down, the positron eventually annihilates into two 511 keV photons.

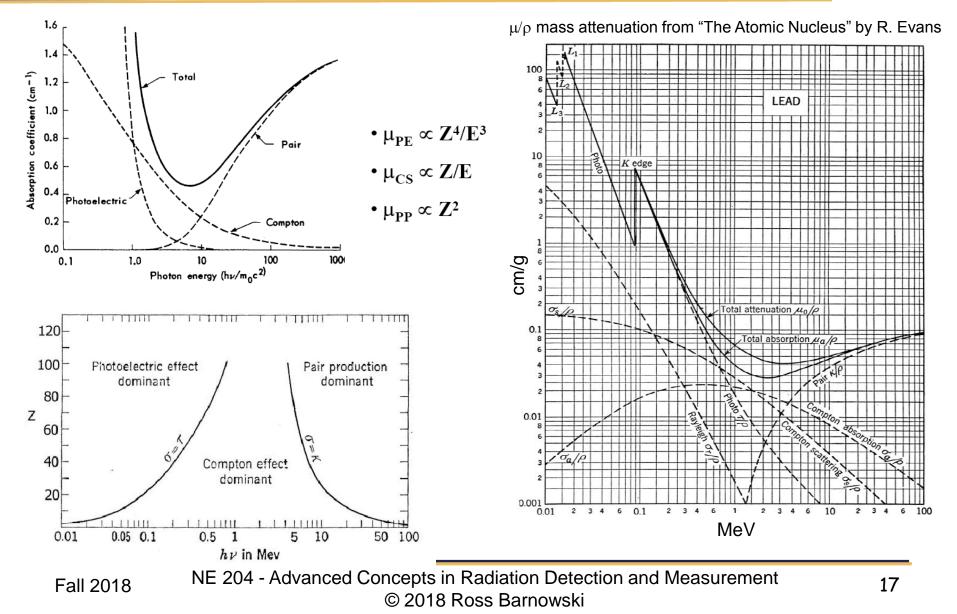
$$E_{e^{-}} + E_{e^{+}} = E_{\gamma} - 2m_e c^2$$

$$\sigma_{PP} \propto Z^2 \ln(E_{\gamma} - 2m_e c^2)$$

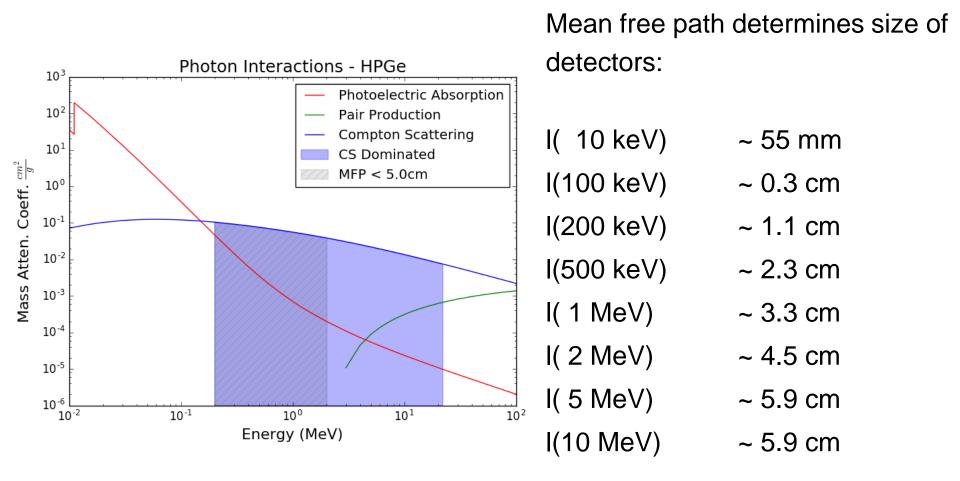


Absorption of Gamma Rays









Interactions of Neutrons



- A beam of neutrons passes through material until each undergoes a collision at random and is removed from beam (strong interaction...)
 - In contrast to photons, the neutrons are "scattered" by nuclei and usually leave only a portion of their energy in the medium until they are very slow and can get absorbed.
 - Intensity drops as well as the neutron energy continuously.
 - The degradation of the beam intensity follows Beer-Lampert exponential attenuation law:

$$I = I_0 e^{-\mu x}, \quad \mu = \mu_{scattering} + \mu_{(n,\gamma)} + \dots$$

- We have to distinguish several classes of interactions:
 - Elastic scattering (n,n)
 - Inelastic scattering (n,n')
 - Radiative capture (n,γ)
 - Charged-particle production reaction (n,p), (n, α),...
 - Fission ²³⁵U, ²³⁹Pu,...(n,f)



- $\sigma_{\text{total}} = \sigma_{s} + \sigma_{a}$ (cross section σ expressed in barns [10⁻²⁴ cm²]) $- \sigma_{s} = \sigma_{e} + \sigma_{i}$
- $-\sigma_a = \sigma_\gamma + \sigma_f + \sigma_p + \sigma_\alpha + \dots$ (n,γ) + elastic + (n,p), (n,α) in (n,2n), (n,3n), very light elements + (n,fission) (n,pn) (n,p2n), (laogarithmic scale) in fissile isotopes etc. +inelastic potential scattering + fission in heavy $(\sigma = constant)$ elements σ∝1/E region $\sigma_{total} = \sigma_s$ 9 MeV 8 MeV 10 Cross-section resonance 5500 barns ! region Cross-section (barns) from =1eV $\sim 1/v \sim E^{-1/2}$ (higher in light and near-magic nuclei) ³He(n,p) $B(n, \alpha)$ ⁶Li(n,α) 0.1 10- 10^{-10} 10 10^{2} 10^{3} 10⁵ 1(4 1(6 1(7 Energy (eV)

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